

STCE Newsletter

16 Feb 2015 - 22 Feb 2015



Published by the STCE - this issue : 27 Feb 2015. Available online at <http://www.stce.be/newsletter/> .

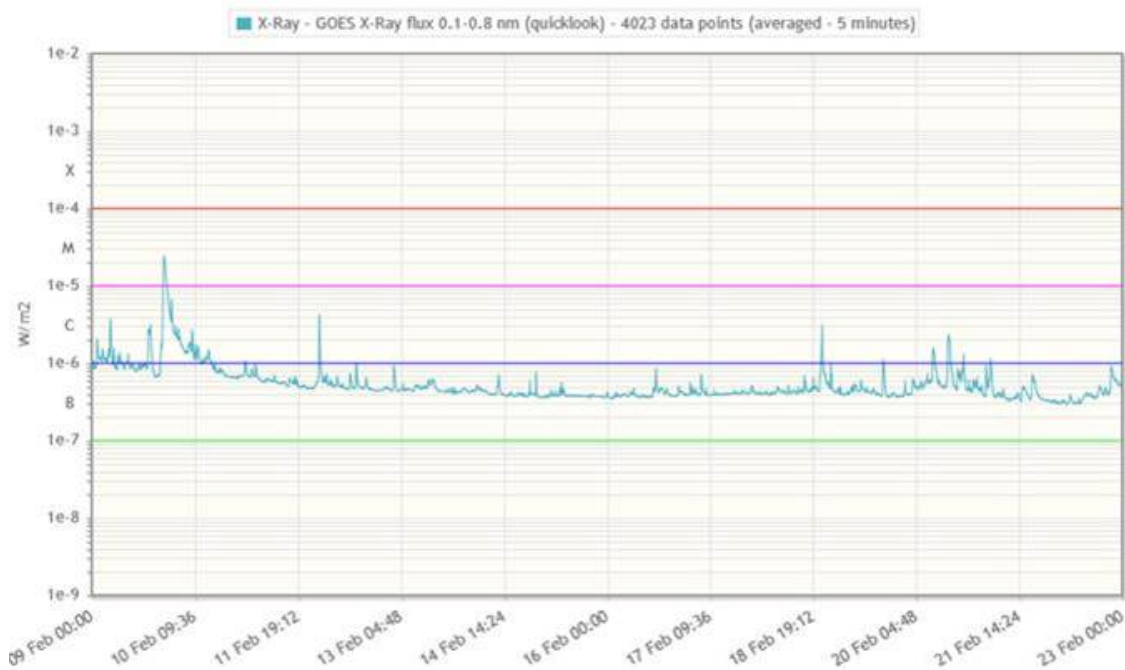
The Solar-Terrestrial Centre of Excellence (STCE) is a collaborative network of the Belgian Institute for Space Aeronomy, the Royal Observatory of Belgium and the Royal Meteorological Institute of Belgium.

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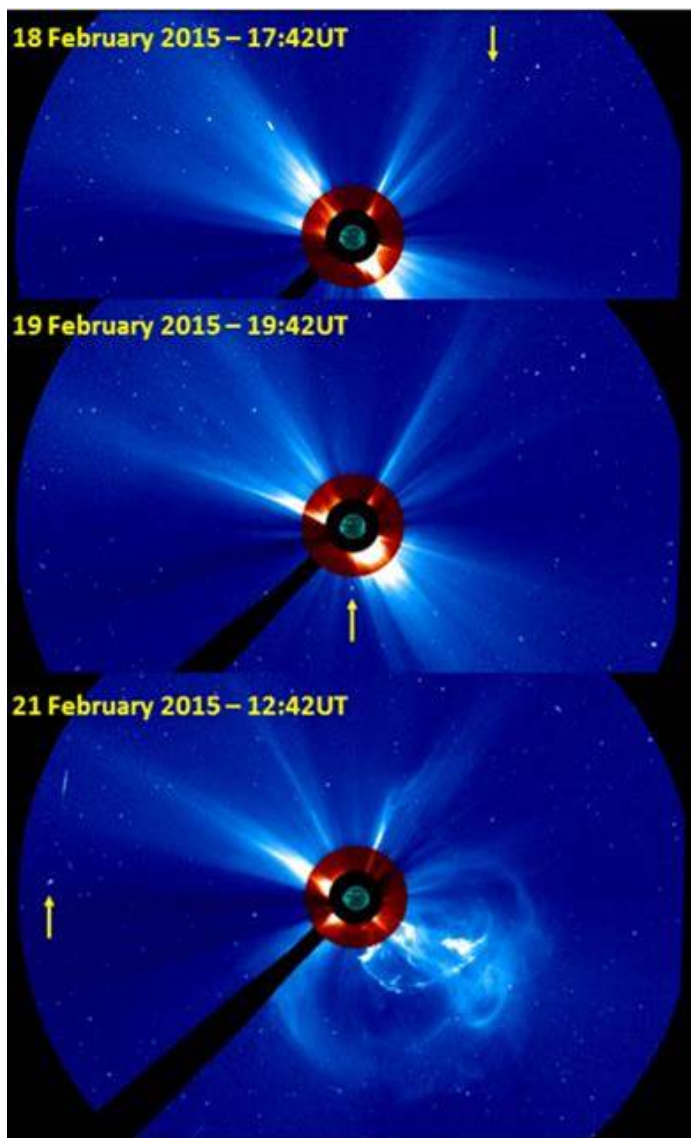
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1. An icy visitor from outer space

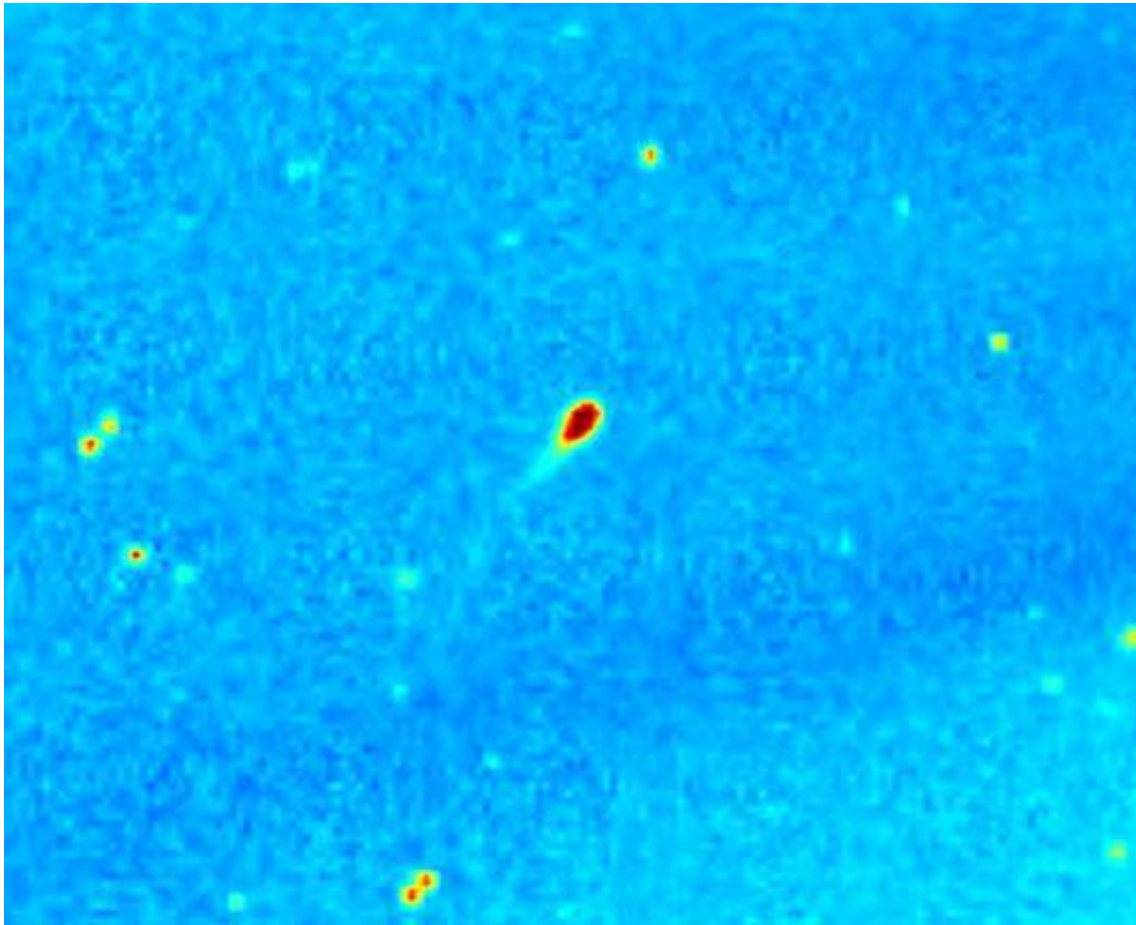
After the M2 flare on 9 February, the Sun decided to take a break, with only low-level C-class flares during the subsequent two weeks. In fact, for 5 consecutive days (from 13 till 17 February), not even a C-class flare was registered, as can be seen in graph underneath.



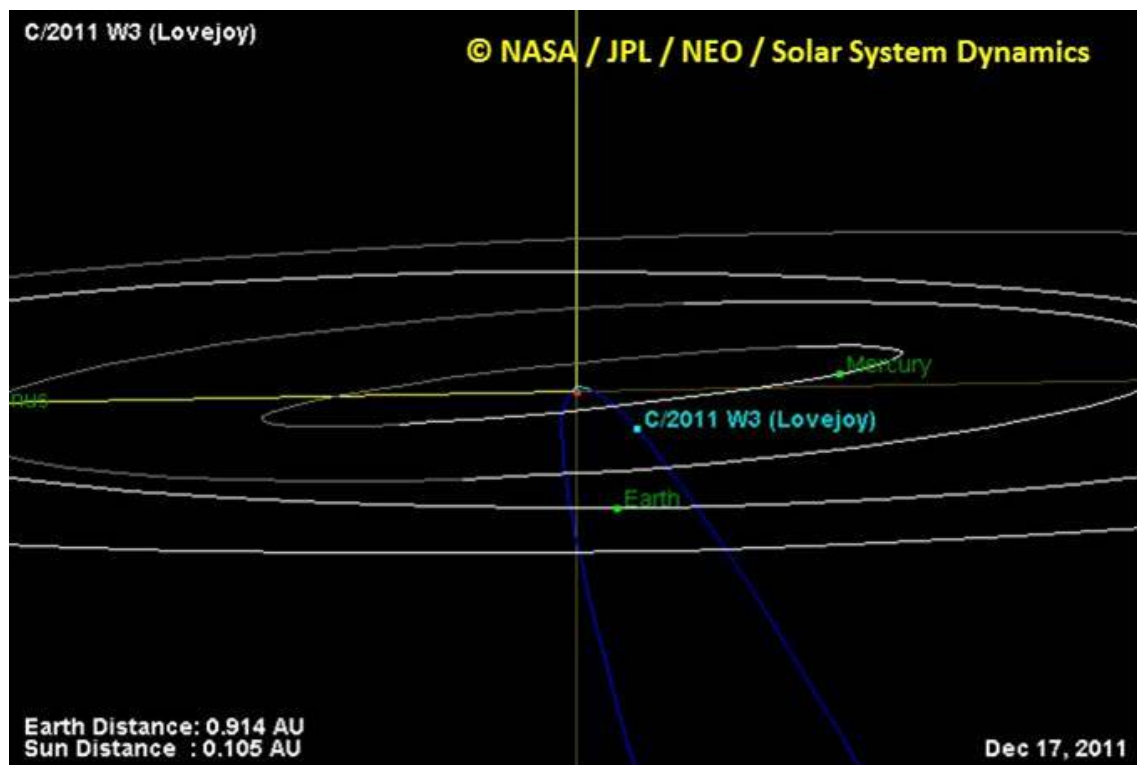
So, there seemed to be very little of interest to write about, but fortunately the folks at SOHO alerted the community of a tiny comet passing through the field of view of SOHO's coronagraphs, see their pick of the week at <http://sohowww.nascom.nasa.gov/pickoftheweek/old/20feb2015/>). A coronagraph is an instrument that blocks out the solar disk so that the immediate vicinity around the Sun becomes visible, similar to what our Moon does during a total solar eclipse. Though SOHO's coronagraph is used to observe coronal mass ejections (CME), it has also made SOHO the most successful comet hunter so far. Indeed, the newly found dot is the 2875th comet on SOHO's very long list of comet discoveries. Quite a few of those were actually made by amateur astronomers who were examining SOHO's imagery.



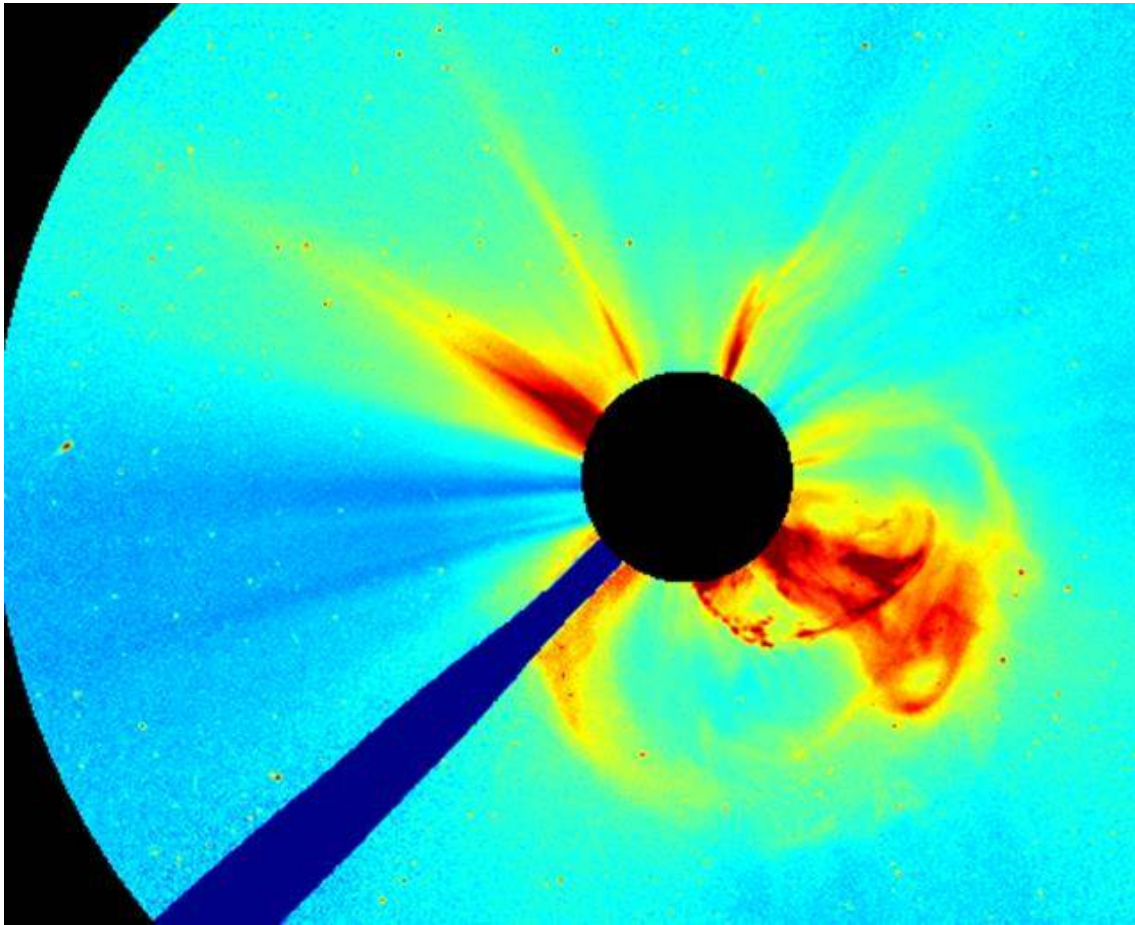
As the comet SOHO-2875 was approaching the Sun, it appeared as a fast-moving star-like object on 18 February (see image above on top). It could hardly be distinguished from the starry background, except for its motion of course. Scientists were not expecting much of this ball of ice and dust, until the comet all of a sudden started to display a tiny tail after its swing-by of the Sun. Astronomers are now eagerly awaiting its appearance in the evening sky to see if it can be observed by telescope.



What seems already certain is that -unlike most of the comets discovered by SOHO- this one does not belong to the Kreutz group. This is a group of comets that all follow pretty much the same orbit that brings them very close to the Sun ("sungrazers"). It is believed that they are all fragments of a comet that broke apart several centuries ago. Most of the Kreutz members do not survive their kiss and ride with the Sun, comet Lovejoy (2011) being a notable exception (see sketch of orbit below). That SOHO-2875 does not belong to the Kreutz group is apparent from its motion. The Kreutz members all pass most of their orbit below the ecliptic (the plane created by the yearly movement of the Earth around the Sun), which make them also rather insensitive to gravitational perturbations from the planets. Only during their close encounter with the Sun do they briefly stay above the ecliptic. The completely different orbits can be seen in this movie at <http://youtu.be/YvIs-jEiv2c> where the motion of SOHO-2875 is compared to that of comet Lovejoy.



SOHO-2875 seems also to have been just in time to avoid being hit by 2 CMEs. In particular the second CME on 21 February was quite strong and speedy (average plane-of-the-sky speed of about 900 km/s). It originated from the backside of the Sun and concerned the eruption of the very long filament that was visible from Earth just 2 weeks ago, as discussed in this news item at <http://stce.be/news/294/welcome.html> The image underneath (as well as in the movie) is a false color clip showing the comet to the left and the CME from the filament eruption expanding to the lower right. The CME did not have an Earth directed component.



Credits - Images for the movies were taken from SOHO (<http://sohowww.nascom.nasa.gov/home.html>), SDO (<http://sdo.gsfc.nasa.gov/data/aiahmi/>), NASA/JPL/NEO/Solar System Dynamics (<http://neo.jpl.nasa.gov/orbits/>), and (J)Helioviewer (<http://helioviewer.org/>).

2. PROBA2 Observations (16 Feb 2015 - 22 Feb 2015)

Solar Activity

Solar flare activity fluctuated between very low and low during the week.

In order to view the activity of this week in more detail, we suggest to go to the following website from which all the daily (normal and difference) movies can be accessed:<http://proba2.oma.be/ssa>

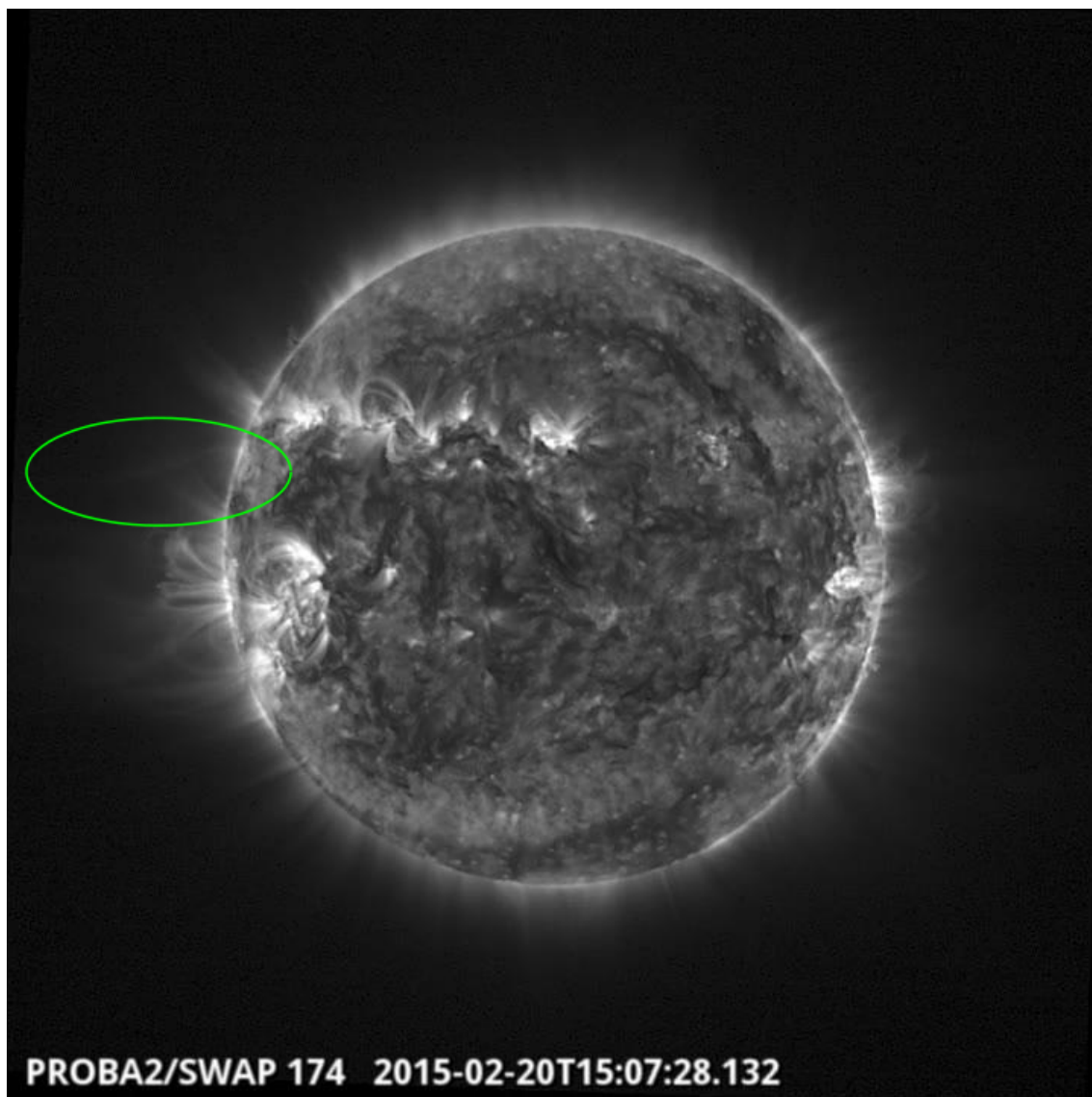
This page also lists the recorded flaring events.

A weekly overview movie can be found here (SWAP week 256).

http://proba2.oma.be/swap/data/mpg/movies/weekly_movies/weekly_movie_2015_02_16.mp4

Details about some of this week's events, can be found further below.

Friday Feb 20

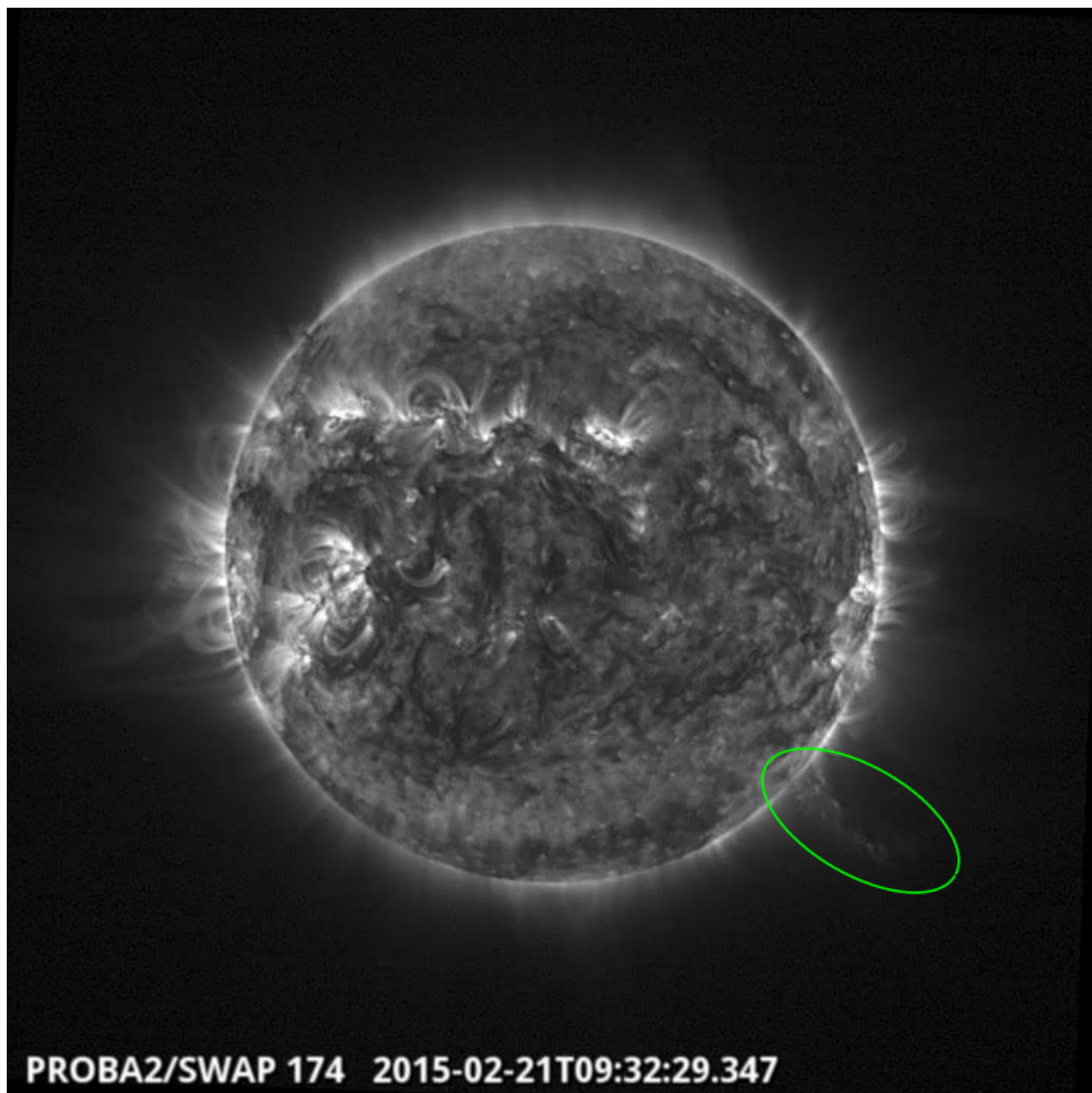


Saturday Feb 21

Eruption on the east limb @ 15:07 SWAP image

Find a movie of the event here (SWAP movie)

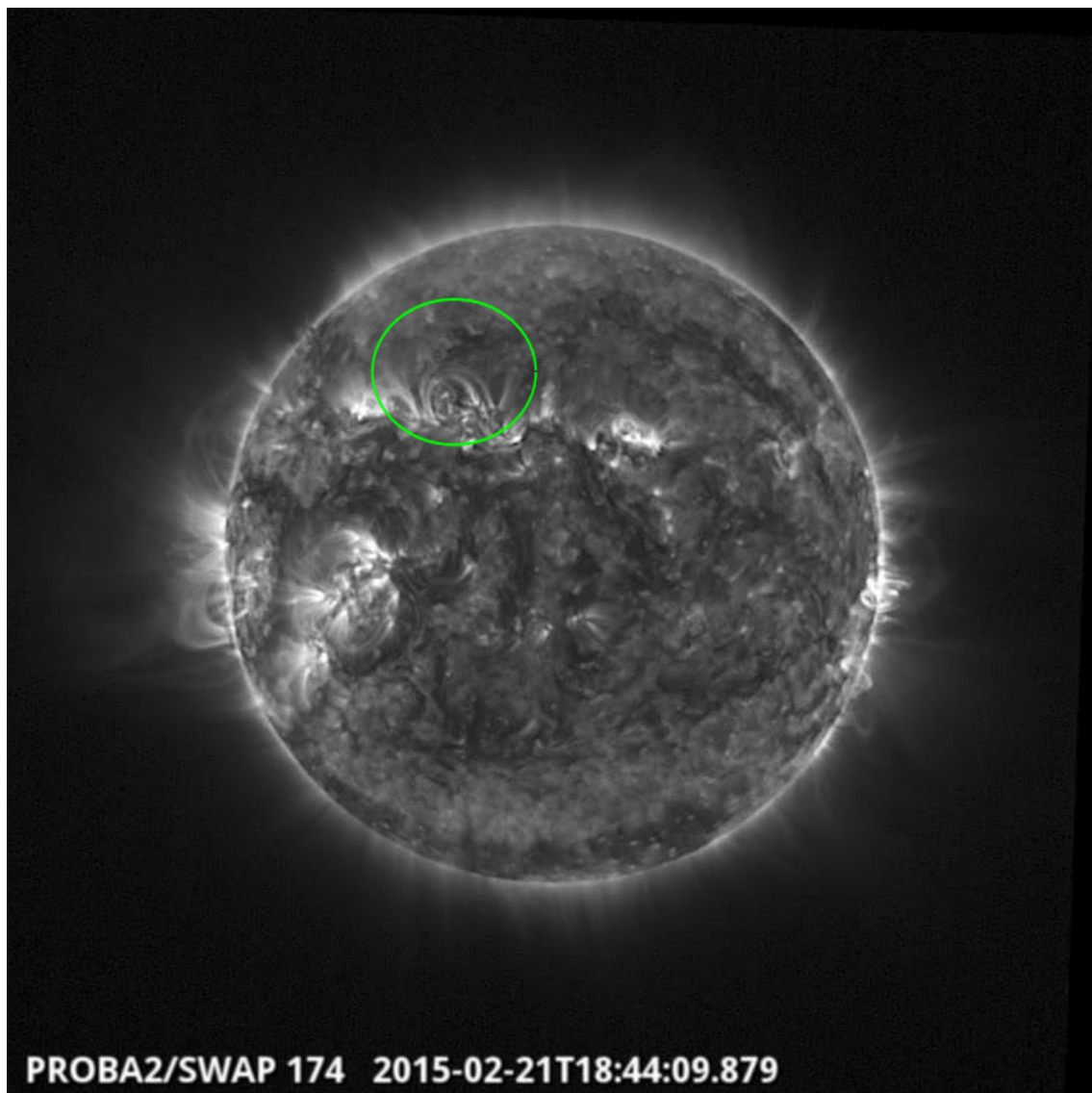
http://proba2.oma.be/swap/data/mpg/movies/20150220_swap_movie.mp4



Eruption on the west limb @ 09:32 SWAP image

Find a movie of the event here (SWAP movie)

http://proba2.oma.be/swap/data/mpg/movies/20150221_swap_movie.mp4



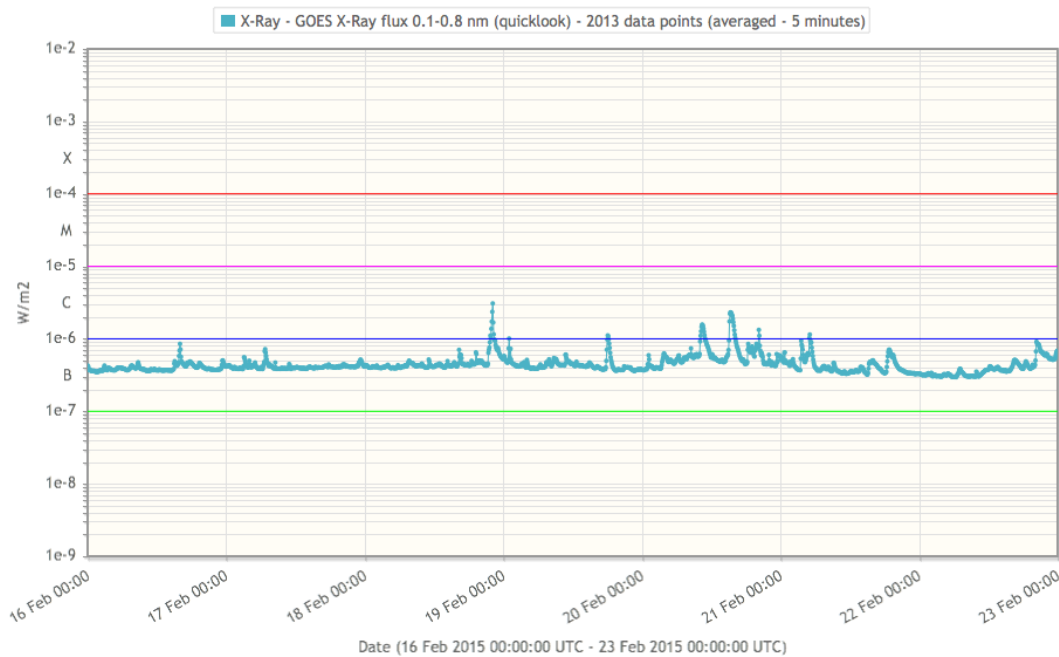
Eruption on the west limb @ 18:44 SWAP image

Find a movie of the event here (SWAP movie)

http://proba2.oma.be/swap/data/mpg/movies/20150221_swap_movie.mp4

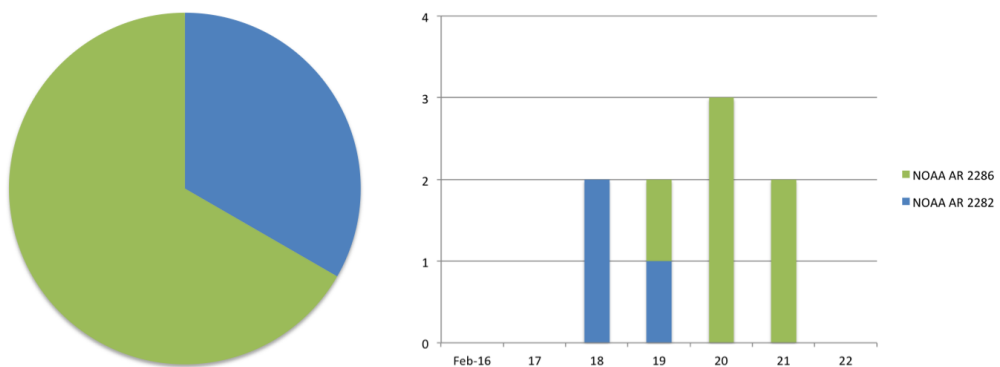
3. Review of solar activity

Solar activity was very low to low throughout the week. The background level of the X-ray flux was at B-level as can be seen in the graph below where the GOES X-ray flux 0.1-0.8 nm is plotted from Feb 16 to 22.



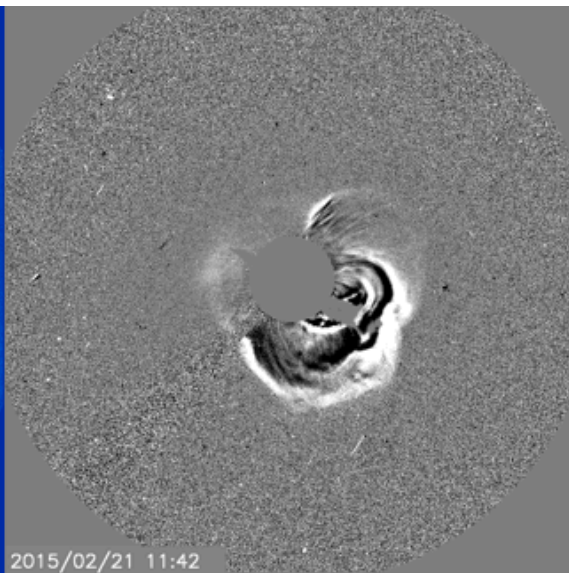
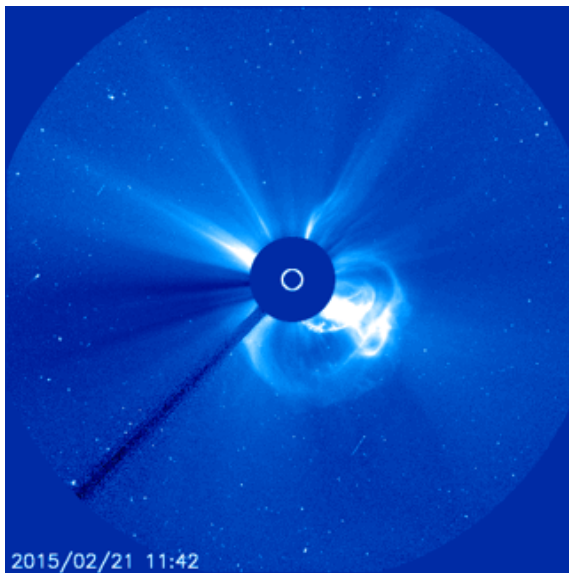
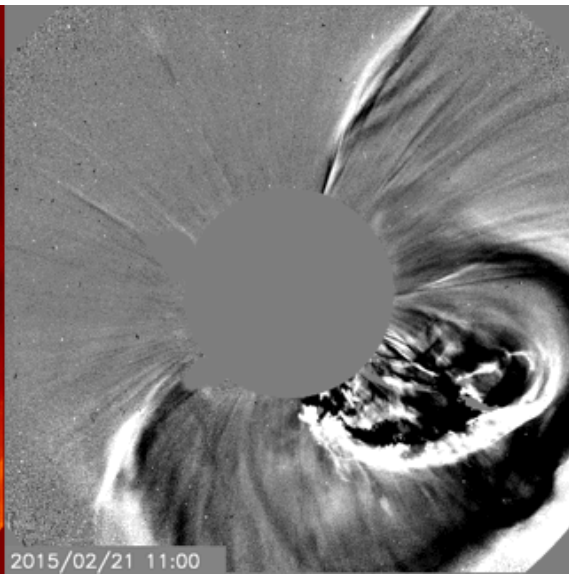
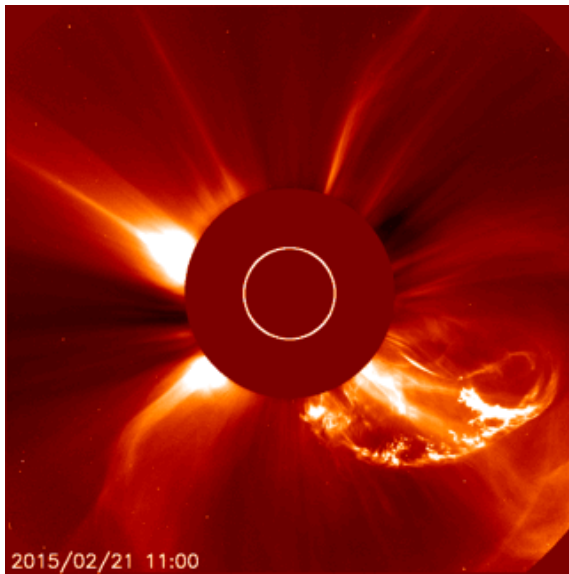
Only nine lower C-class flares and a bunch of B-class flares were produced. The largest activity originated from NOAA active regions (ARs) 2282 and 2286, with a C2.3 flare as strongest Xray event (peak at 15:22 UT on February 20). NOAA ARs 2287, 2289 and 2290 only produced some tiny flaring.

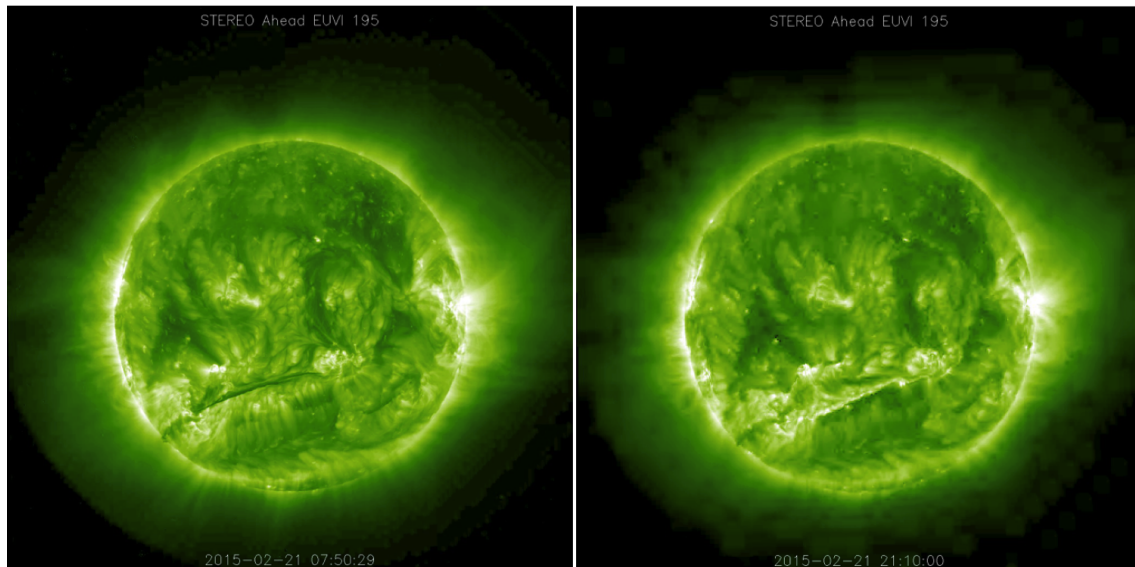
Distribution of C flares, Feb 16 – 22, 2015



The left chart gives an overview of the total number of flares per NOAA AR region for the indicated week. The right chart gives an overview of the flaring activity per NOAA AR per day. No M- or X-flares occurred.

An asymmetric halo CME was detected in coronagraphic imagery (first measurement in SOHO LASCO/C2 at 9:24 UT on February 21), associated to a backside eruption of a large filament in the southern hemisphere that was present for more than one and a half of a solar rotation.

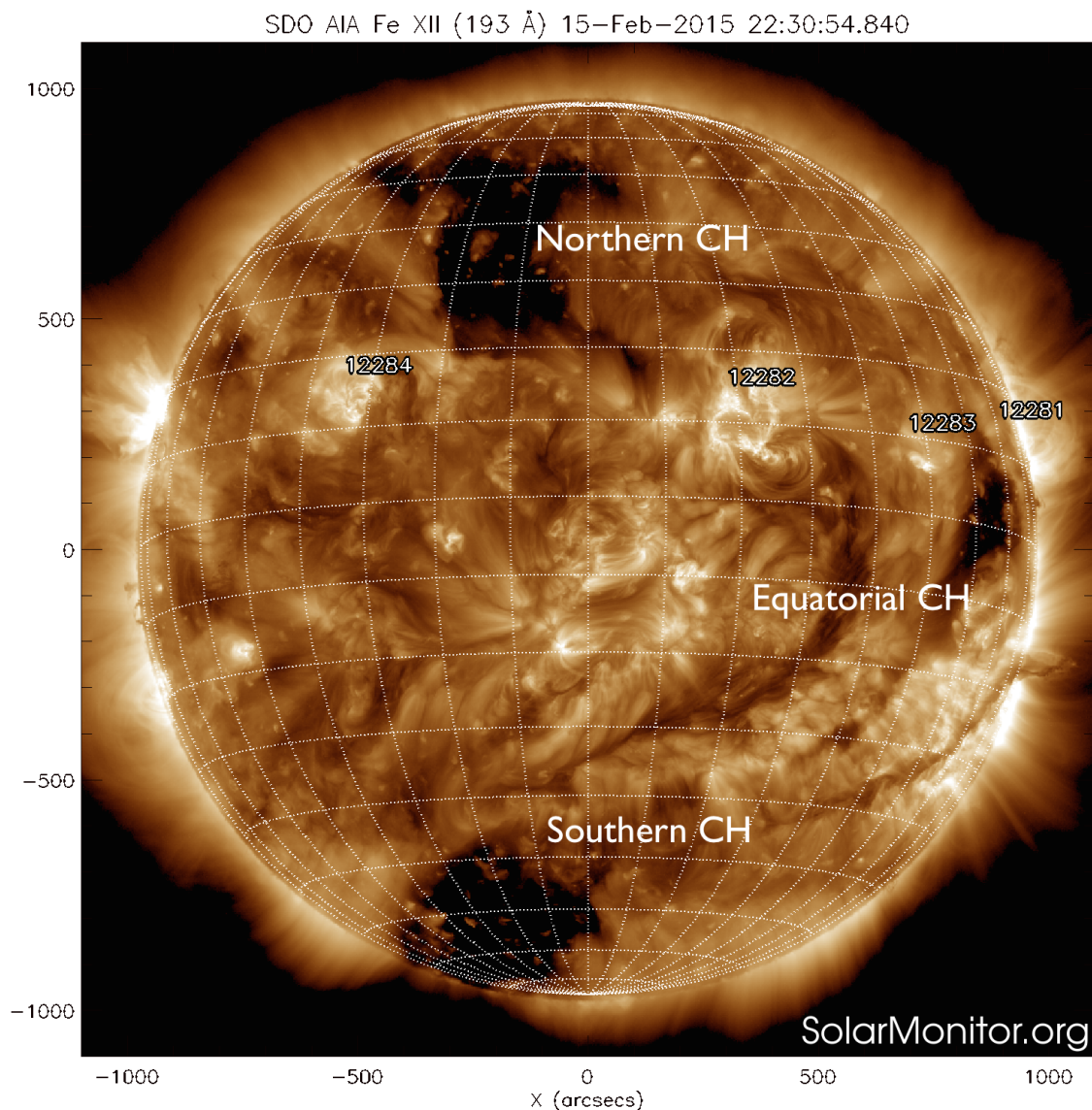




The red and blue image are taken by the coronagraph LASCO onboard SOHO. The grey images are difference images and show what changed between the two subtracted images. Difference images help to see changes. The green images are taken by the STEREO A spacecraft and show the back side of the Sun. You see the large erupting filament (left) and the post-eruption loops (right) .

The CME was mainly heading to the southwest with an estimated projected speed of approximately 1000 km/s. The proton flux (at > 10 MeV energy levels) has risen to enhanced levels starting around 12 UT, but did not pass the event threshold.

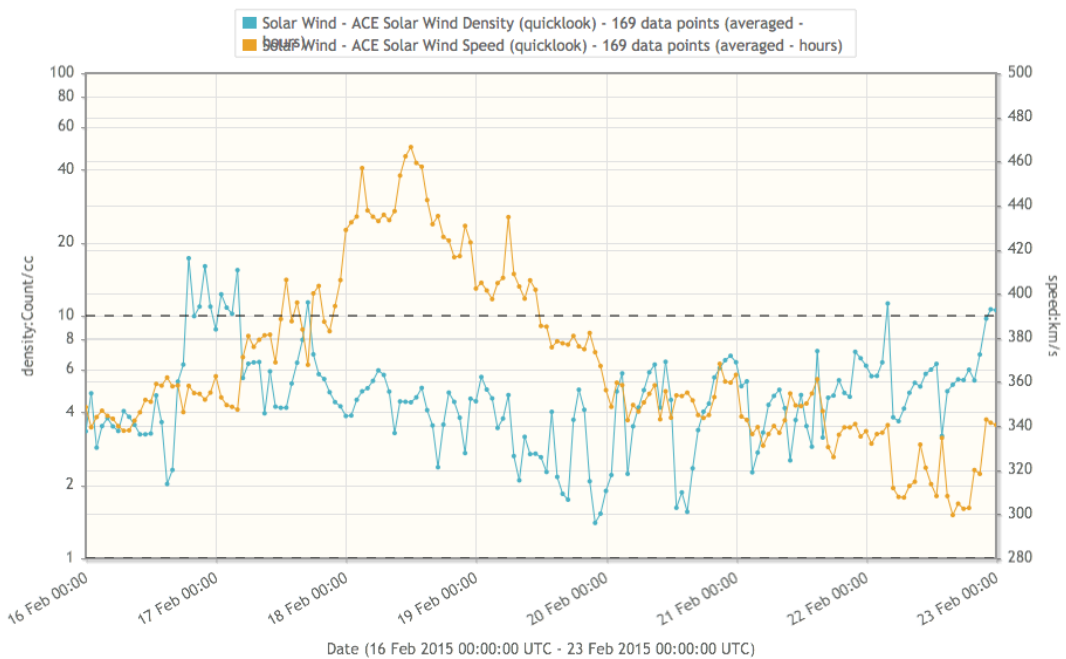
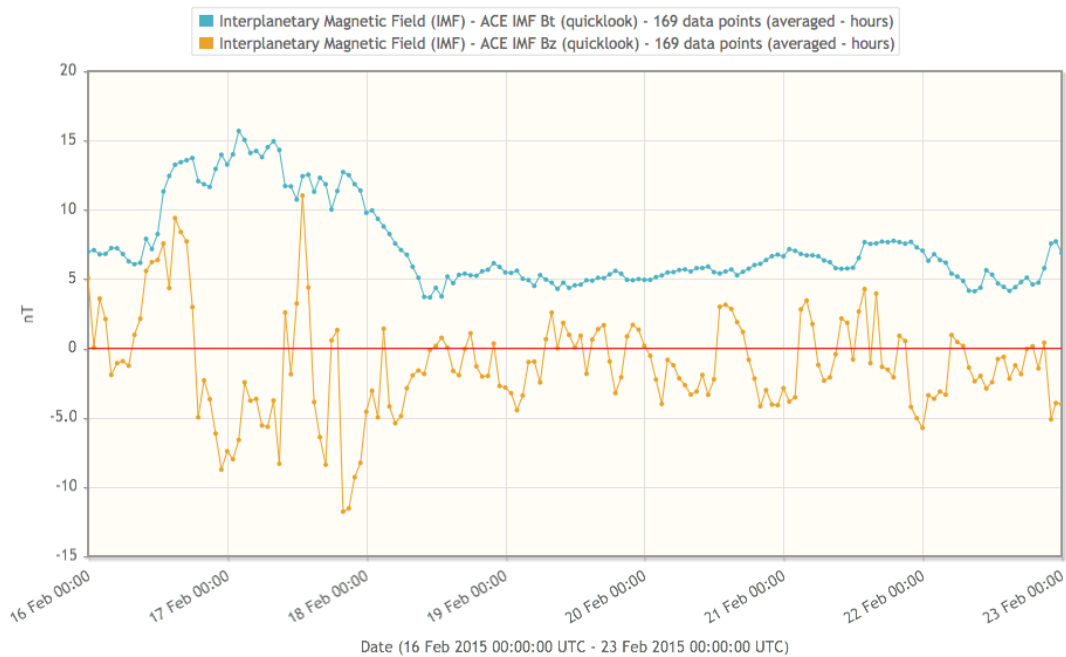
A filament eruption was visible in EUV imagery starting in PROBA2/SWAP at 18:09 UT on February 21, accompanied by a long-duration B7 flare.



A few coronal holes were observed in SDO imagery: an equatorial negative polarity coronal hole (CH) located near the west limb on February 16, a southern one of negative polarity and a northern one of positive polarity.

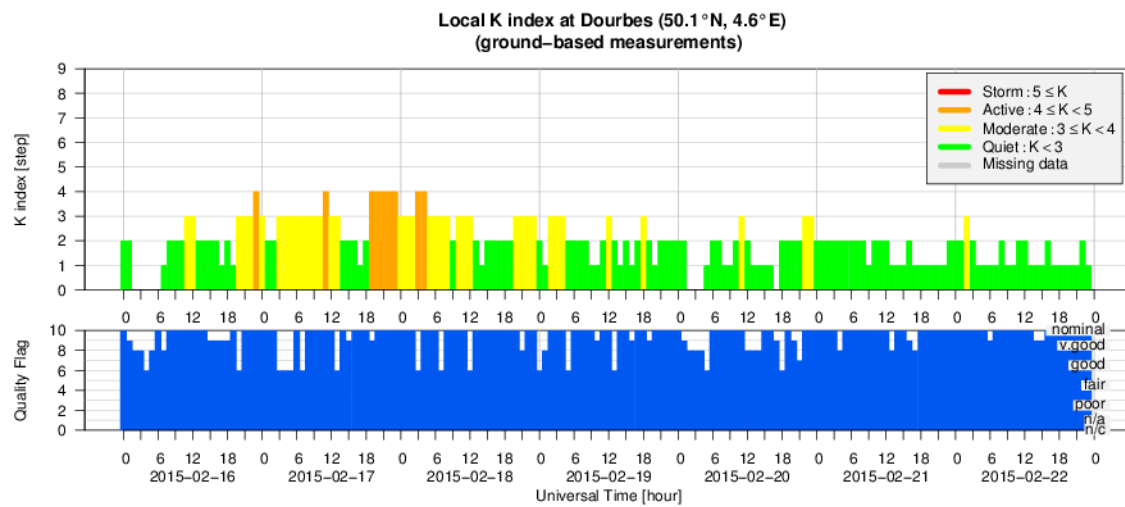
4. Review of geomagnetic activity (16 Feb 2015 - 22 Feb 2015)

For most of the week Earth was facing nominal solar wind conditions with an interplanetary magnetic field magnitude below 10 nT. During the period between 12:00 UT on February 16 and 8:00 UT on February 17, the IMF magnitude was slightly enhanced to maximally 16 nT. Solar wind speed was mostly between 300 and 400 km/s, except for the period from early February 17 till the second half of February 19 on during which a maximum of about 480 km/s was reached.



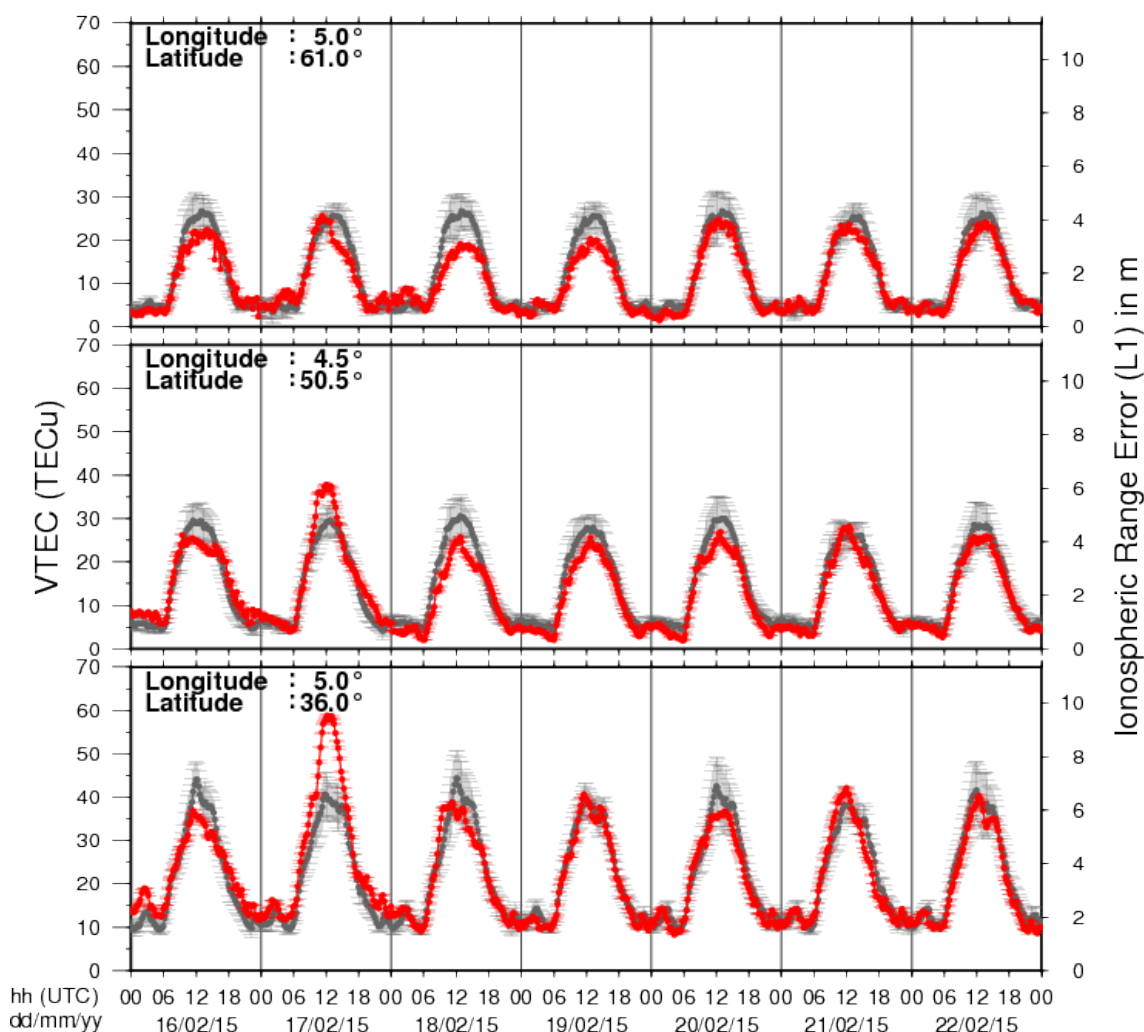
The disturbed solar wind conditions (magnetic field and solar wind speed) starting on February 16 are associated to the influence of a compression region followed by a solar wind stream from the equatorial negative polarity CH. Geomagnetic conditions were unsettled to active ($K=3$ to 4) during several time slots from 18 UT on February 16 to 12 UT on February 18. Quiet to unsettled geomagnetic conditions were observed during the remainder of the week.

5. Geomagnetic Observations at Dourbes (16 Feb 2015 - 22 Feb 2015)



6. Review of ionospheric activity (16 Feb 2015 - 22 Feb 2015)

VTEC Time Series



The figure shows the time evolution of the Vertical Total Electron Content (VTEC) (in red) during the last week at three locations:

- a) in the northern part of Europe(N61°, 5°E)
- b) above Brussels(N50.5°, 4.5°E)
- c) in the southern part of Europe(N36°, 5°E)

This figure also shows (in grey) the normal ionospheric behaviour expected based on the median VTEC from the 15 previous days.

The VTEC is expressed in TECu (with $\text{TECu} = 10^{16}$ electrons per square meter) and is directly related to the signal propagation delay due to the ionosphere (in figure: delay on GPS L1 frequency).

The Sun's radiation ionizes the Earth's upper atmosphere, the ionosphere, located from about 60km to 1000km above the Earth's surface. The ionization process in the ionosphere produces ions and free electrons. These electrons perturb the propagation of the GNSS (Global Navigation Satellite System) signals by inducing a so-called ionospheric delay.

See http://stce.be/newsletter/GNSS_final.pdf for some more explanations ; for detailed information, see http://gnss.be/ionosphere_tutorial.php

7. Future Events

For more details, see <http://www.spaceweather.eu/en/event/future>

Conference on Sun-Climate Connections (SCC 2015) in Kiel, Germany

Start : 2015-03-16 - End : 2015-03-19

This international conference will provide an overview of our current understanding of Sun-Climate Connections starting at processes on the Sun itself over space weather and solar wind towards solar influence on the upper atmosphere down to the ocean. It will also provide insights into the heatedly debated role of the Sun in climate change. In four sessions the various contributions of solar variability influence on Earth's climate will be presented and discussed by bringing together solar physicists, space scientists, atmospheric scientists, climate modellers, and paleoclimatologists.

We expect contributions from scientists participating in SCOSTEP/ROSMIC, SPARC-SOLARIS/HEPPA, the EU cost network TOSCA, as well as any other interested scientists. The conference will last three full days, beginning Monday morning, 16 March 2013. The programme will consist of invited and keynote lectures, a few contributed oral presentations and ample time dedicated to poster sessions. The fourth day will be devoted to public outreach activities as well as panel discussions.

Website: <http://scc.geomar.de/>

URSI AT-RASC 2015 in Gran Canaria, Spain

Start : 2015-05-18 - End : 2015-05-22

URSI AT-RASC 2015 will be the first edition of the newly established triennial URSI Atlantic Radio Science Conference as one of the URSI Flagship Conferences. AT-RASC 2015 will have an open scientific program composed of submitted papers within the domains covered by all ten Commissions of URSI.

Website: <http://www.at-rasc.com/>

MHD waves and instabilities in the solar atmosphere in Budapest, Hungary

Start : 2015-05-25 - End : 2015-05-29

25-27 May 2015: BUKS 2015 - MHD waves: Observational aspects from ground to space - MHD waves: Theory - where are we? - MHD instabilities

27-28 May 2015: Ruderman Honorary meeting - Theory of linear MHD waves - MHD waves instabilities - Non-linear waves in plasmas

29 May 2015: Joint BUKS/Ruderman's conferences excursion - Boat excursion to Szentendre, Visegrad and Esztergom

Website:

http://swat.group.shef.ac.uk/Conferences/BUKS_2015/index.html

Los Alamos Space Weather Summer School, in Los Alamos, NM, USA

Start : 2015-06-01 - End : 2015-07-24

The Space Weather Summer School at Los Alamos National Laboratory, established in 2011 under the founding Director Josef Koller, is dedicated to space weather, space science and applications. Every year we solicit applications for the Los Alamos Space Weather Summer School. This summer school is sponsored and supported by a number of organizations at LANL. This year our top sponsors include the Los Alamos Institute of Geophysics, Planetary Physics and Signatures (IGPPS) and the Laboratory Directed Research and Development Office (LDRD). The summer school brings together top space science students with internationally recognized researchers at LANL in an educational and collaborative atmosphere.

Website:

<http://www.swx-school.lanl.gov/>

Solar dynamo frontier workshop in Boulder, CO (USA)

Start : 2015-06-09 - End : 2015-06-12

The last five years have seen substantial progress in our understanding of the solar dynamo, fueled by continuing advances in observations and modeling. With the launch of NASA's Solar Dynamics Observatory (SDO) in 2010 came an unprecedented window on the evolving magnetic topology of the Sun, highlighting its intricate 3D structure and global connectivity. The Helioseismic Magnetic Imager (HMI) instrument on SDO in particular has provided potentially transformative yet enigmatic insights into the internal dynamics of the solar convection zone that underlie the dynamo. Attempts to detect subsurface convective motions from helioseismic inversions have yielded only upper limits on the large-scale convective amplitude, challenging our understanding of global solar convection. Yet, potential signatures of giant cells have been detected in photospheric Dopplergrams. Estimates of the meridional flow from HMI and complementary instruments (SOHO/MDI and GONG) have been equally tantalizing and enigmatic. Several disparate techniques, including local and global helioseismic inversions and correlation tracking of surface features, have yielded evidence of a multi-cellular meridional flow but they differ on the detailed flow structure and amplitude. This multi-cellular meridional flow has potentially profound implications for flux-transport dynamo models that previously assumed a very different structure with a single circulation cell per hemisphere.

Website:

<https://www2.hao.ucar.edu/Workshop/Solar-Dynamo-Frontiers>

CISM Space Weather Summer School in Boulder, CO, USA

Start : 2015-07-13 - End : 2015-07-24

The CISM Summer School is intended to give students a comprehensive immersion in the subject of space weather: what it is, what it does, and what can be done about it. Space weather is many things: beautiful when seen through the eyes of a sun-viewing telescope, fascinating when studied for its alien worlds of magnetic structures and phenomena, awesome when witnessed as a solar eruption or auroral storm, and devastating to the users of services it disrupts. Space weather links the Sun, the Earth, and the space in between in a branching chain of consequences. Weather systems on the Sun can spawn interplanetary storms of colossal size and energy that envelop the whole planet in electrical hurricanes. Such storms attack high-tech, complex, and expensive technological systems that provide much of the infrastructure that allows modern society to function.

Website:

<https://www2.hao.ucar.edu/Events/2015-CISM-Summer-School>

Loops7: Heating of the Magnetically Closed Corona in Cambridge, UK

Start : 2015-07-21 - End : 2015-07-23

The conference will review past and recent achievements, as well as future challenges in the field of solar coronal loop physics.

Website:

<http://www.damtp.cam.ac.uk/user/astro/cl7/index.html>

Heliophysics Summer School 2015: Seasons in Space: Cycles of variability of Sun-Planet systems, in Boulder, CO, USA

Start : 2015-07-28 - End : 2015-08-04

Heliophysics is all of the science common to the field of the Sun-Earth connections. This fast-developing field of research covers many traditional sub-disciplines of space physics, astrophysics, and climate studies. The NASA Living with a Star program, with its focus on the basic science underlying all aspects of space weather, acts as a catalyst to bring the many research disciplines together to deepen our understanding of the system of systems formed by the Sun-Earth connection.

Website:

<http://www.heliophysics.ucar.edu/>

34th International Cosmic Ray Conference (ICRC) in The Hague, The Netherlands

Start : 2015-07-30 - End : 2015-08-06

The 34th International Cosmic Ray Conference (ICRC) will be held from July 30 to August 6, 2015, in The Hague, The Netherlands. It is an important and large conference in the field of Astroparticle Physics. The ICRC covers: cosmic-ray physics, solar and heliospheric physics, gamma-ray astronomy, neutrino astronomy, and dark matter physics.

Website: <http://icrc2015.nl>

SOLARNET III / HELAS VII: The Sun, the stars, and solar-stellar relations, in Freiburg (Germany)

Start : 2015-08-31 - End : 2015-09-04

The purpose of this conference is to discuss the latest questions and results in solar and stellar physics. Solar and stellar seismology will be one particular focus but contributions on all aspects of solar-stellar relations will be welcome. We aim to establish links and synergies between the day- and night-time fields of astrophysics.

Website:

<http://www.iac.es/congreso/solarnet-3meeting/>

1st Joint Solar Probe Plus-Solar Orbiter Workshop, in Florence (Italy)

Start : 2015-09-02 - End : 2015-09-04

The Workshop will address how the joint exploration of the corona and inner heliosphere will lead to advances in our understanding of coronal heating and solar wind acceleration, the magnetic and plasma structure of the heliosphere, and the acceleration of energetic particles at shocks and flares. The workshop will inspire research that will make use of SO and SPP observations within the context of the NASA Heliophysics Observatory System and identify key areas for preparatory research. Synergistic observations from other ground based and space based assets will also be addressed.

Website:

<http://www.solarprobeplus.org/2015/>

Ground-based Solar Observations in the Space Instrumentation Era in Coimbra, Portugal

Start : 2015-10-05 - End : 2015-10-09

This CSPM-2015 scientific meeting will cover various aspects of solar dynamic and magnetic phenomena which are observed over the entire electromagnetic spectrum: white-light, Hα, Ca II, and radio from ground and in a variety of other wavelengths (white light, UV and EUV, and X-rays) from space. Emphasis will also be placed on instrumentation, observing techniques, and solar image processing techniques, as well as theory and modelling through detailed radiative transfer in increasingly realistic MHD models. The long-term (cyclic) evolution of solar magnetism and its consequence for the solar atmosphere, eruptive phenomena, solar irradiation variations, and space weather, will be in focus. Here, special attention will be devoted to the long-term observations made in Coimbra and also to the results of the SPRING / SOLARNET and SCOSTEP VarSITI studies. In particular, the weak solar activity during the current solar maximum will be discussed. Finally, since this meeting is organised around the 90th anniversary of performing the first spectroheliographic observations in Coimbra, a session will be specially dedicated to new solar instruments (both ground-based and space-borne) that will give access to unexplored solar atmospheric features and dynamic phenomena over the coming years.

Website:

<http://www.mat.uc.pt/~cspm2015/>

2015 Sun-Climate Symposium in Savannah, Georgia, USA

Start : 2015-11-10 - End : 2015-11-13

Observations of the Sun and Earth from space have revolutionized our view and understanding about impacts of solar variability and anthropogenic forcing on Earth climate. For more than three solar cycles since 1978, the total and spectral solar irradiance (TSI and SSI) and global terrestrial atmosphere/surface have been observed continuously, enabling unprecedented quality data for Sun-climate studies. The primary objective of this symposium is to convene climate scientists, solar physicists, and experimentalists together for a better understanding how Earth climate system changes and responds to solar variability.

Website: <http://lasp.colorado.edu/home/sorce/news-events/meetings/2015-sun-climate-symposium/>

41st COSPAR Scientific Assembly in Istanbul, Turkey

Start : 2016-07-30 - End : 2016-08-07

The 41st COSPAR Scientific Assembly will be held in Istanbul, Turkey from 30 July - 7 August 2016. This Assembly is open to all bona fide scientists.

Website:

<https://www.cospas-assembly.org/>

8. New documents in the European Space Weather Portal Repository

See <http://www.spaceweather.eu/en/repository>

eHEROES - De Zon

Theoretical course on the Sun and space weather for participants to the astronomy course in Public Observatory MIRA, Grimbergen (Belgium). Given on 19 March 2014 for 35 attendees.

<http://www.spaceweather.eu/en/repository/show?id=557>

STCE - Space weather science, infrastructure, services and products: SW events and impact

Presentation given during a users' visit about the STCE operational space weather services and products.

<http://www.spaceweather.eu/en/repository/show?id=558>

STCE - Space weather science, infrastructure, services and products: Service Centers

Presentation given during a users' visit about the STCE operational space weather services and products.

<http://www.spaceweather.eu/en/repository/show?id=559>

STCE - Space weather science, infrastructure, services and products: Operational Software and Products

Presentation given during a users' visit about the STCE operational space weather services and products.

<http://www.spaceweather.eu/en/repository/show?id=560>

STCE - Space weather science, infrastructure, services and products: SW forecast

Presentation given during a users' visit about the STCE operational space weather services and products.

<http://www.spaceweather.eu/en/repository/show?id=561>

STCE - Space weather science, infrastructure, services and products: SEP

Presentation given during a users' visit about the STCE operational space weather services and products.

<http://www.spaceweather.eu/en/repository/show?id=562>

STCE - Space weather science, infrastructure, services and products: Ionospheric event

Presentation given during a users' visit about the STCE operational space weather services and products.

<http://www.spaceweather.eu/en/repository/show?id=563>

eHEROES - De verschillende vormen van zonneactiviteit en hun invloed op de mens en zijn technologie

Invited review submitted to the journal Revue E. This article is the first in a series of 3 articles. De Zon, Helios, Sol, ... er bestaan vele namen voor die gele bol die dagelijks ons hemelgewelf doorkruist en onze warmte- en lichtbron bij uitstek is. Dankzij satellietwaarnemingen hebben we onze ster leren kennen als een dynamisch en explosief hemelobject dat aan de basis ligt van het zogenaamde ruimteweer dat een belangrijke impact heeft op onze technologie.

<http://www.spaceweather.eu/en/repository/show?id=564>

eHEROES - Onderzoek naar de zonnecorona

On the occasion of the solar eclipse of March 20, 2015, we contributed to the March 2015 edition of the amateur astronomer journal 'Zenit'. We highlighted the outcome of LASCO onboard of SOHO and focussed on the role of the STCE on space missions like PROBA2, PROBA3 and Solar Orbiter. The text is written in Dutch.

<http://www.spaceweather.eu/en/repository/show?id=565>